

- 1-i A vector has components $A^a = \{0, -2, 3, 5\}$ in the frame K . What are the components of the vector in a frame which moves with the velocity $0.8c$ in the positive z direction relative to K .
- 1-ii Two events in special relativity are separated by a space-like interval, $ds^2 < 0$. Is there an inertial frame where the two events are simultaneous?
- 1-iii In special relativity: argue that a free electron can neither absorb nor emit a photon.
- 1-iv Consider a 3-dimensional Euclidean space with Cartesian coordinates. Is there a difference between vectors with index up and index down?
- 1-v Consider a vector A^a . Is the four-component object $\{\frac{1}{A^0}, \frac{1}{A^1}, \frac{1}{A^2}, \frac{1}{A^3}\}$ a vector?
- 2-i In Schwarzschild coordinates $\{t, r\}$ the half-life of an elementary particle measured at rest at large distance from the center is Δt . What is its half-life at rest at a distance r from the center? What is the proper half-life?
- 2-ii Find out how the expression $A^a_{;b}$ transforms under a change of coordinates. Does it obey the tensor transformation law?
- 2-iii Show that the Christoffel symbol Γ^a_{bc} is symmetric under exchange of the lower indices.
- 2-iv Consider a scalar function of coordinates $\phi(x)$. Find out whether the objects $\frac{\partial \phi}{\partial x^a}$ and $\frac{\partial^2 \phi}{\partial x^a \partial x^b}$ are tensors in special and general relativity.
- 2-v Prove that in a locally inertial frame
- $$R_{abcd,e} = \frac{1}{2}(g_{ad,bce} - g_{ac,bde} + g_{bc,ade} - g_{bd,ace).$$
- 2-vi Starting with Minkowski metric $ds^2 = \eta_{ab}dx^a dx^b$, show that the coordinate transformation $r = \sqrt{x^2 + y^2 + z^2}$, $\theta = \arccos(z/r)$, $\varphi = \arctan(y/x)$ leads to metric $ds^2 = dt^2 - dr^2 - r^2(d\theta^2 + \sin^2 \theta d\varphi^2)$.
- 3-i Assume that the components of the metric tensor g_{ab} do not depend on the coordinate x^1 . Show that for a free moving body the component u_1 of the body's four-velocity is then conserved.
- 3-ii Show that Maxwell equations $F^b_{;a} = 4\pi j^a$ it follows that $j^a_{;a} = 0$. Hint: prove first that $\sqrt{-g}A^c_{;c} = (\sqrt{-g}A^c)^{;c}$.
- 3-iii Use the definition of the invariant volume element (called covariant in lecture notes) to calculate the proper area of a sphere.
- 3-iv Show that for an equatorial orbit in the Schwarzschild metric the quantity u_φ^2 is conserved (where u^φ is the φ -component of the four-velocity, $u^a = dx^a/ds$).

- 4-i In the weak field limit show that $g_{ab} = \eta_{ab} + h_{ab}$ where $h_{yz} = A \sin \omega(t-x)$, $h_{tt} = 2f(t-x)$, $h_{tx} = -f(t-x)$, all other $h_{ab} = 0$, f is an arbitrary function, satisfies the linearized Einstein equation in vacuum. Hint: see the gravitational waves chapter.
- 4-ii Let us define the generalized force \mathcal{F}_a acting on a particle with mass m through the equation of motion $mDu_a/ds = \mathcal{F}_a$. Recall that from $u^a u_a = 1$ it follows that $Du_a u^a = 0$ and therefore $\mathcal{F}_a u^a = 0$. Show that the electromagnetic force satisfies this condition. Show that a force in the form $\mathcal{F}_a = -\partial V/\partial x^a$, where V is a scalar function of coordinates, generally does not satisfy this condition. What would a discovery of such force mean?
- 4-iii Consider a nearly circular orbit of a planet around a star in Newtonian mechanics and in General Relativity. Derive the equation for a small radial perturbation of the orbit and find its angular period. Relate to the post-Newtonian perihelion precession.
- 4-iv Consider a flat radiation-dominated Friedman universe. Show that at early times it expands as $a \propto \sqrt{t}$.